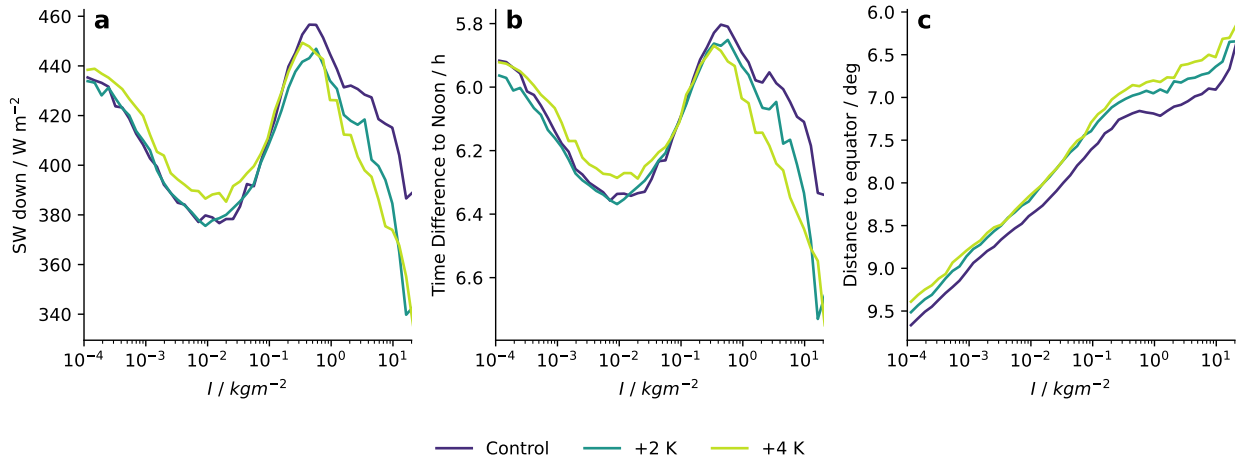




15 **S1 Influence of Time of Occurrence on High-Cloud Radiative Effect**

16 The SW  $C(I)$  becomes stronger for more incoming solar radiation, which can be modulated by the local  
17 time at which the respective cloud occurs, and its distance to the equator. To first order, the incoming  
18 SW radiation is governed by the timing of the clouds. This becomes evident by the solar radiation closely  
19 following the time difference to noon (Fig. S1).



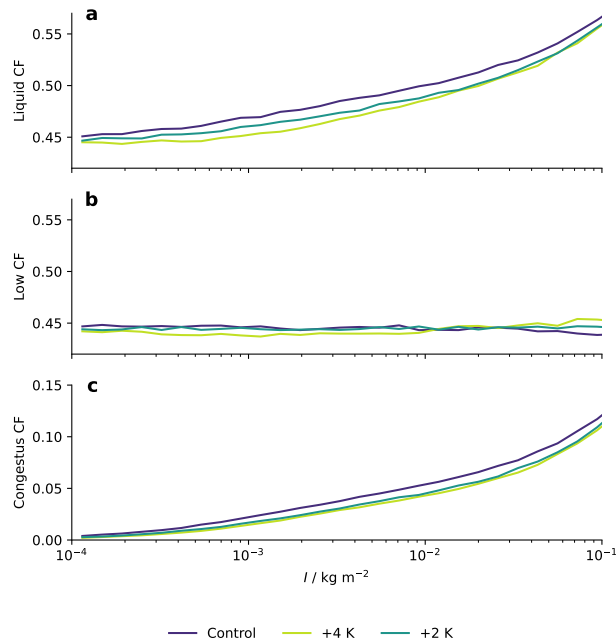
**Fig. S1** (a) Incoming SW radiation, (b) time difference to noon and (c) distance to the equator binned by ice water path ( $I$ ).

## 20 S2 Congestus Feedback

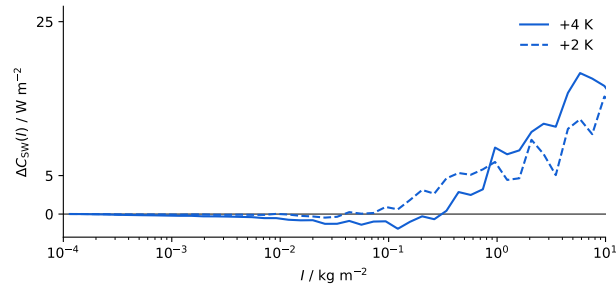
21 Liquid clouds can either be part of a high cloud if the frozen and liquid condensate sufficiently overlap, or not  
 22 (Eq. 3). At low  $I$ , overlap between frozen and liquid clouds is rare, but it exists. We refer to such overlapping  
 23 clouds as congestus clouds, since they typically contain more liquid than frozen condensate. The rest of the  
 24 liquid clouds are referred to as low clouds.

25 Liquid clouds reduce as surface temperature increases (Fig. S2a), which causes a reduction in both  
 26 low clouds (Fig. S2b) and congestus clouds (Fig. S2c). Congestus clouds also carry less liquid as surface  
 27 temperatures increase. The median liquid water path of congestus clouds at  $I < 10^{-1} \text{ kg m}^{-2}$  is  $0.169 \text{ kg m}^{-2}$   
 28 (control),  $0.152 \text{ kg m}^{-2}$  (+2 K) and  $0.145 \text{ kg m}^{-2}$  (+4 K).

29 The reduction of congestus clouds together with their reduced amount of liquid condensate causes the SW  
 30  $C(I)$  to become less negative at  $I < 10^{-1} \text{ kg m}^{-2}$  in response to surface warming (Fig. 1d). This becomes  
 31 apparent when we calculate  $\Delta C_{\text{SW}}(I)$  only for frozen clouds, i.e. liquid condensate is not included as part  
 32 of the high clouds. In this case, the positive  $\Delta C_{\text{SW}}(I)$  at  $I < 10^{-1} \text{ kg m}^{-2}$  does not exist (Fig. S3).



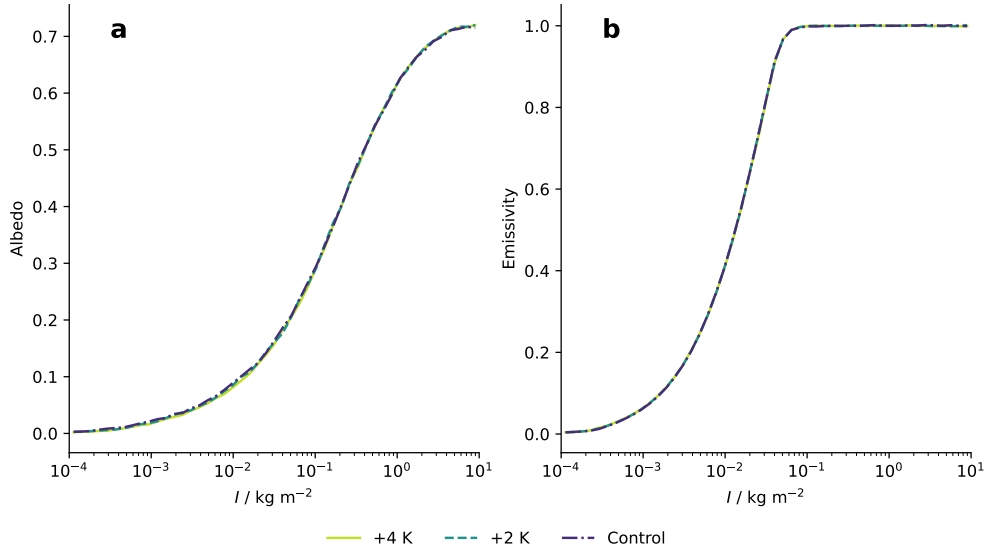
**Fig. S2** (a) Fraction of all liquid clouds, (b) fraction of low clouds (liquid clouds that are not part of the high clouds) and (c) fraction of congestus clouds (liquid clouds that are part of the high clouds).



**Fig. S3** Change in the SW  $C(I)$  for frozen clouds only.

### 33 S3 Albedo and Emissivity of High Clouds

34 We test whether changes in the microphysics of high clouds in response to surface warming alter the high-  
35 cloud opacity by calculating the high-cloud albedo and emissivity [1]. Both the albedo and the emissivity  
36 of high clouds remain remarkably constant despite surface warming (Fig. S4). We conclude from this that  
37 potential changes in the microphysics of high clouds in response to surface warming do not alter their opacity  
38 in our simulations.



**Fig. S4** (a) Albedo of high clouds and (b) emissivity of high clouds.

### 39 S4 Constant Ozone Simulation

40 The temperature profile of the +4 K warming simulation with prescribed constant ozone simulation shows  
41 clear tropopause warming. This is not the case for warming simulations with interactive ozone, for which  
42 the temperature profile shifts upward.

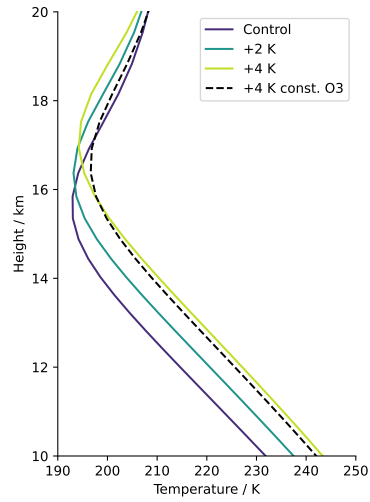
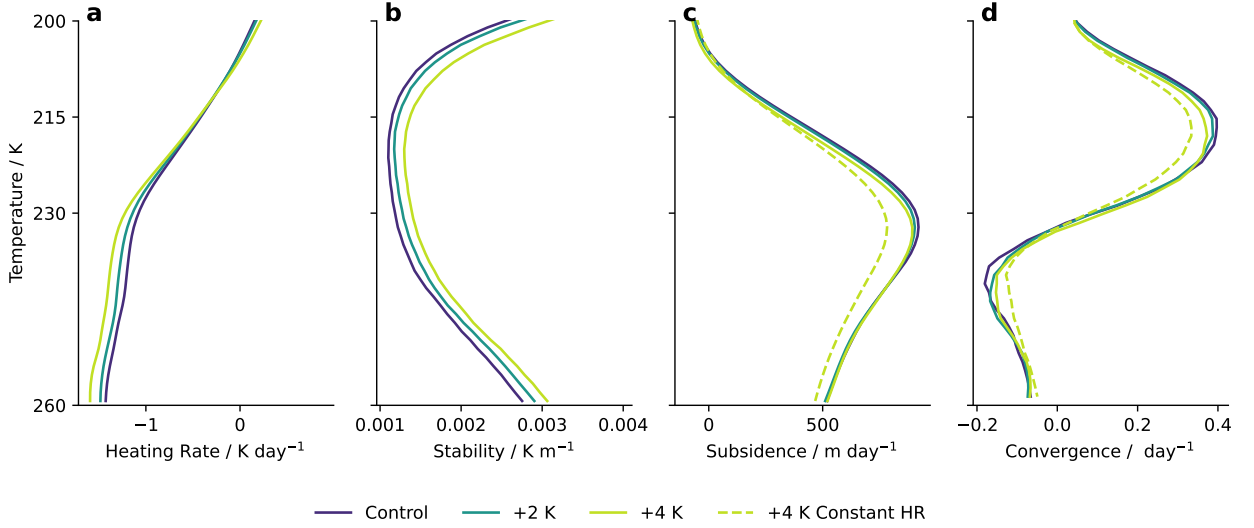


Fig. S5 Mean clear-sky temperature profiles.

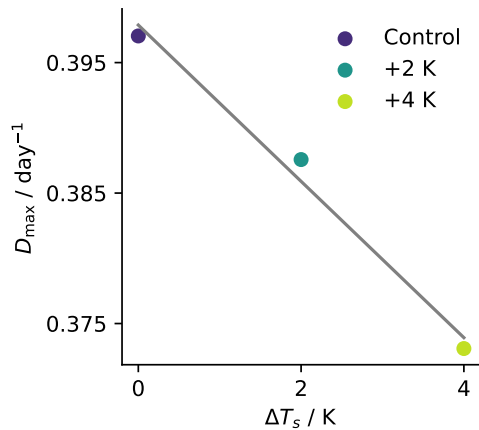
## 43 S5 Stability Iris

44 The static stability is increasing with surface warming as expected (Fig. S6b); however, this is partly cancelled  
 45 by an increase in clear-sky radiative cooling (Fig. S6a). Therefore, changes in the subsidence profiles and the  
 46 convergence are small (Fig. S6c, d). This cancellation was also found by previous studies, but seems more  
 47 pronounced in our simulations [2].



**Fig. S6** (a) Mean profiles within the clear-sky atmosphere of heating rate, (b) static stability, (c) radiatively driven subsidence and (d) horizontal convergence of radiatively driven subsidence. The dotted line in c and d shows results for the +4 K simulation with the heating rate fixed at the control profile.

48 Regressing the maximum clear-sky convergence ( $D_{\max}$ ) against the surface temperature anomaly yields  
 49 the decrease of  $D_{\max}$  in response to increasing surface temperatures, which is central to the stability iris  
 50 theory (Fig. S7).



**Fig. S7** Maximum clear-sky convergence ( $D_{\max}$ ) plotted against the surface temperature anomaly ( $\Delta T_s$ ) and linear regression of the data (grey line).

## 51 S6 Heating Rates

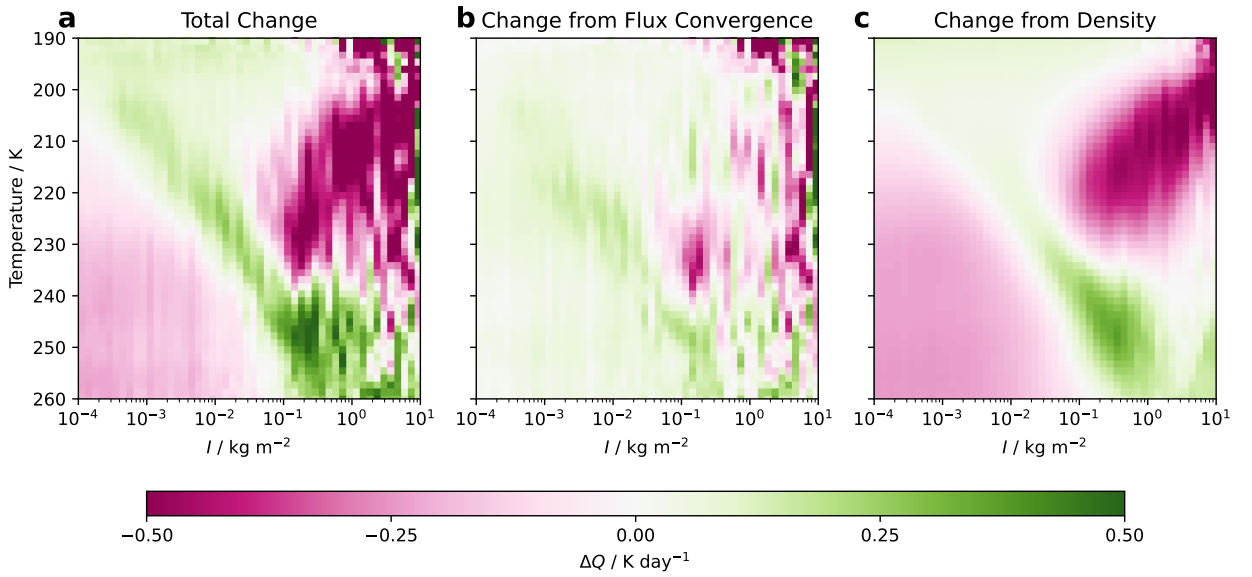
52 Radiative heating rates ( $Q$ ) are calculated as

$$Q = \frac{1}{\rho c_p} \frac{dF}{dz},$$

53 where  $\rho$  is air density,  $c_p$  specific heat capacity of air and  $F$  the radiative flux. As the surface warms, the  
 54 atmosphere expands, and isotherms move to lower pressure and hence lower density. The high clouds follow  
 55 the isotherms [3], and hence also move to lower density. The flux divergence  $\frac{dF}{dz}$  is expected to stay constant  
 56 at a given isotherm, despite the isothermal expansion of the atmosphere [4]. The decrease in density at cloud  
 57 height should therefore imply a higher magnitude of cloud radiative heating [5]. However, changes in the  
 58 diurnal cycle of clouds will affect the SW  $\frac{dF}{dz}$ . We test the relative importance of the density effect and the  
 59 change in flux divergence for the overall change in heating rates by decomposing the change in heating rates:

$$\Delta Q = \underbrace{\Delta \frac{1}{\rho c_p} \frac{dF}{dz}}_{\text{Density}} + \underbrace{\frac{1}{\rho c_p} \Delta \frac{dF}{dz}}_{\text{Flux Divergence}} + \underbrace{\Delta \frac{1}{\rho c_p} \Delta \frac{dF}{dz}}_{\text{non-linear}}.$$

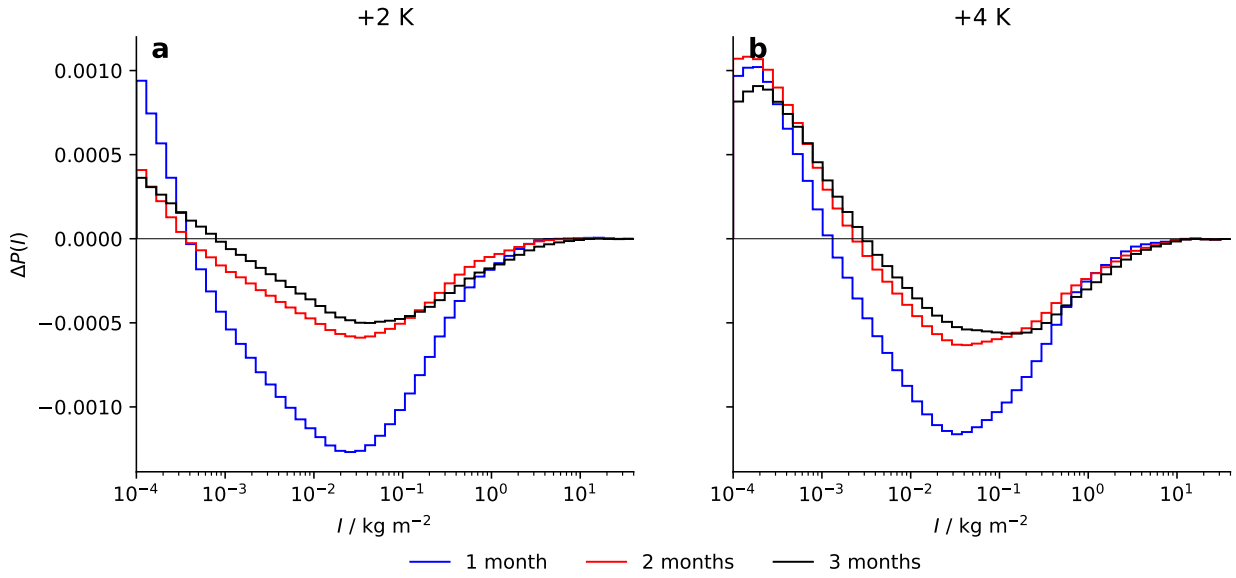
60 This decomposition reveals that the decrease in density is the main driver for changes in heating rates (Fig.  
 61 S8). Changes in the flux divergence exist and cause changes in radiative heating, but they are noisier and  
 62 contribute less to the overall change in radiative heating.



**Fig. S8** (a) Change in cloud radiative heating for the +4 K simulation and the decomposition of the change in heating rate into contributions from (b) the change in flux convergence and (c) the change in density.

## 63 S7 Testing Convergence of Statistics

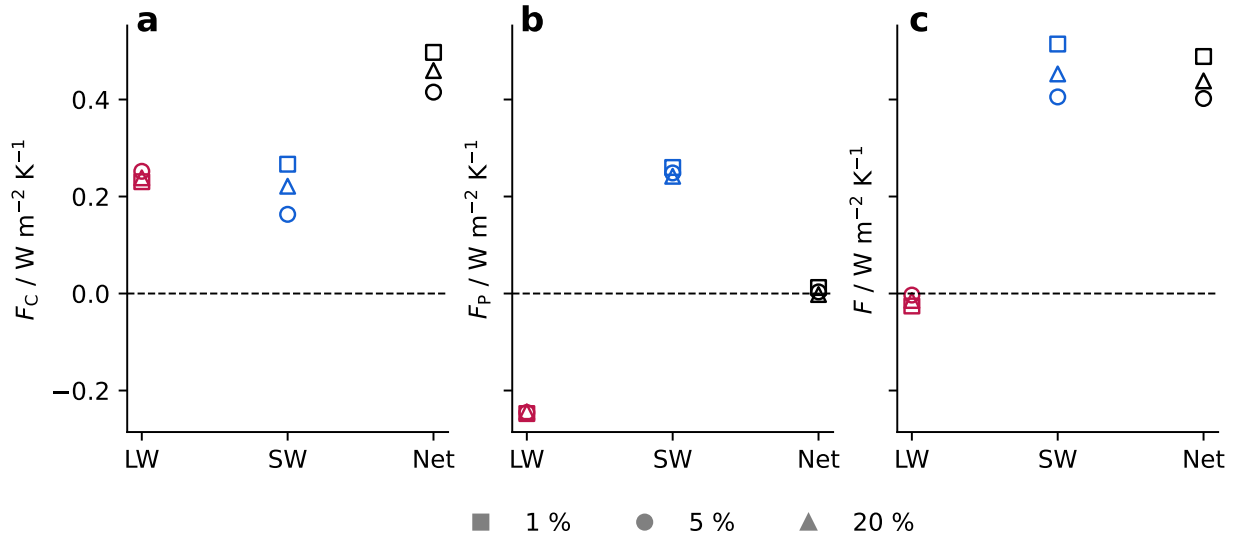
64 To test whether three months of simulation are long enough to reach a statistically stable state, we investigate  
65 the most fragile quantity of our analysis,  $\Delta P$ . When plotted with data from the first month, the first two  
66 months and all three months of simulation,  $\Delta P$  shows convergence (Fig. S9). This proves that three months  
67 of simulation are sufficient to reach a statistically stable state.



**Fig. S9** The difference between the ice water path frequency with respect to the control simulation ( $\Delta P(I)$ ) based on data from the first month, the first two months and all three months of the simulations for (a) the +2 K and (b) the +4 K simulation.

68 **S8 Sensitivity to Connectedness of Liquid and Frozen Clouds**

69 Varying the threshold used to decide whether liquid clouds are part of the high, frozen cloud above mostly  
 70 affects our estimate of the SW  $F_C$  (Fig. S10). This is explained by the SW  $C$  of liquid clouds being much  
 71 stronger than their LW  $C$  [1]. In the paper, we use 5% as the threshold for connectedness.

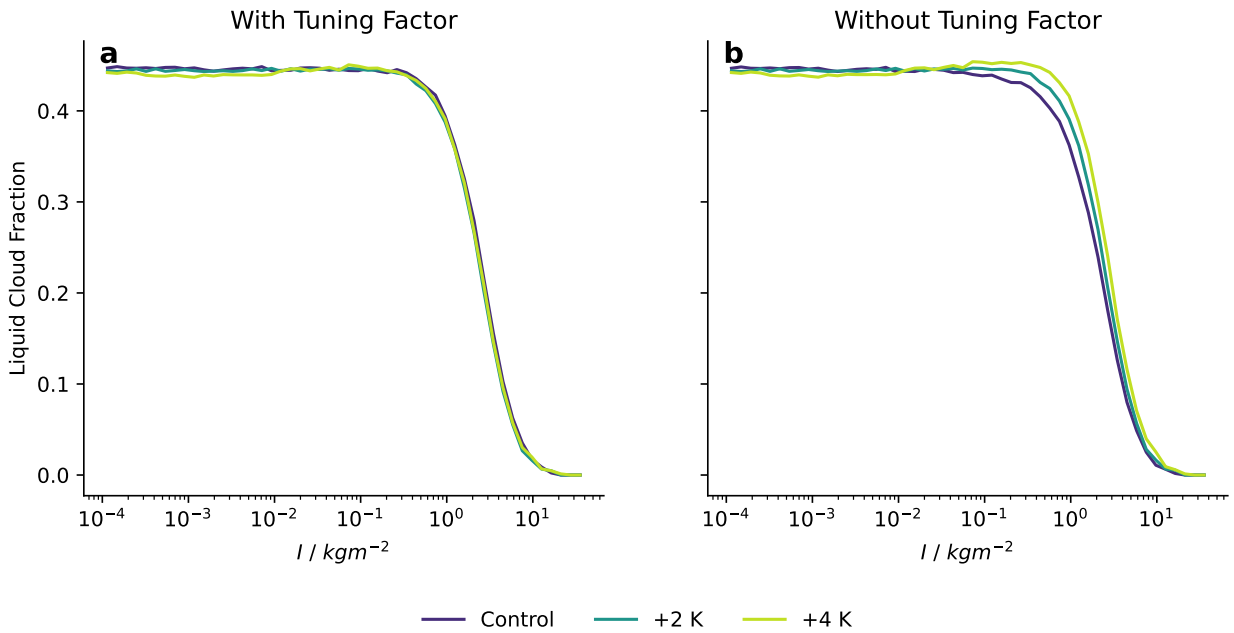


**Fig. S10** High-cloud feedback from the +4 K simulations for different thresholds in the calculation of connectedness. (a) Feedback from changes in  $C(I)$ , (b) feedback from changes in  $P(I)$  and (c) total feedback.

72 **S9 Liquid Clouds at the Transition to Deep Convection**

73 Liquid clouds are typically not connected to the frozen clouds at low  $I$ , but become increasingly connected  
 74 towards high  $I$ . According to our definition of high clouds, liquid clouds are part of the high clouds if  
 75 the liquid and frozen condensate within a column sufficiently overlap (Eq. 3). The transition from liquid  
 76 clouds being unconnected with the frozen clouds to liquid clouds being part of the high cloud occurs at  
 77  $1 \text{ kg m}^{-2} < I < 10 \text{ kg m}^{-2}$  (Fig. S11).

78 Without the tuning factor in Eq. 3, this transition would shift towards higher  $I$  with warming, which  
 79 would make the SW  $C(I)$  become less negative. This is a consequence of liquid clouds being more reflective  
 80 than the ocean surface. We regard this change in the SW  $C(I)$  and the associated feedback to be artificial  
 81 since it depends on subtle details of our definition of high clouds. Therefore, we introduce the tuning factor  
 82 to keep the liquid cloud fraction constant with warming.



**Fig. S11** Fraction of liquid clouds with a liquid water path of more than  $10^{-4} \text{ kg m}^{-2}$  that are not part of high clouds (a) with and (b) without including the tuning factor designed to keep the liquid cloud fraction constant with warming.

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